

Day 16 : Clausius-Mossotti, Boundary conditions, and examples)

Linking the microscopic (often quantum dominated) and macroscopic pictures of physics is one of history's greatest challenges. Stat mech is often about this. Here we'll do it for polarization.

We know  $\vec{P}(r) = n\vec{p}$ , where  $n$  is the number density of atoms

And  $\vec{p} = \alpha \vec{E}$ , where  $\vec{E}$  is all the field acting on the dipole. In a solid this is  $\vec{E}_{\text{Applied}}$  and also the  $\vec{E}$  from nearby atoms.

If we restrict our model to gases, only  $\vec{E}_{\text{Applied}}$  matters and

$$\vec{p} = \alpha \vec{E}_{\text{app}}, \quad \vec{P} = n\alpha \vec{E}_{\text{app}}. \text{ Also, by definition, } \vec{p} = \chi \epsilon_0 \vec{E}$$

We can figure out  $n$  (moles and whatnot), and can measure  $\chi$  with capacitors, but you can't easily measure  $\alpha$  for a single atom or molecule. The above gets us

$$n\alpha \vec{E} = \chi \epsilon_0 \vec{E} = \alpha = \frac{\chi \epsilon_0}{n} \quad (\text{for gases})$$

So we can infer a microscopic parameter from macroscopic ones.

Now, if we don't have a gas, it gets trickier.

$\vec{p}$  is polarized by the fields made by everything but itself, so we say

$$(1) \quad \vec{p} = \alpha(\vec{E}_{\text{total}} - \vec{E}_{\text{self}}) \quad \text{and explicitly subtract out its own field}$$

What is  $\vec{E}_{\text{self}}$ ? Let's approximate the atom as a conducting sphere. We once derived that for such we have

$$\vec{p} = 4\pi a^3 \epsilon_0 \vec{E}_{\text{Applied}} \quad (a \text{ is atomic radius})$$

And the field in a polarized conductor exactly cancels out, so

$$\vec{E}_{\text{self}} = -\vec{E}_{\text{Applied}} = -\frac{\vec{p}}{4\pi a^3 \epsilon_0}$$

Now, in a dense material, the number of atoms per volume is roughly  $1/\text{Volume per atom}$ , so  $n = \frac{1}{4\pi a^3 \epsilon_0}$  and

$$\vec{E}_{\text{self}} = -\frac{n\vec{p}}{3\epsilon_0} = -\frac{\vec{p}}{3\epsilon_0} \quad \text{big } \vec{P}$$

Put this into (1) to get

$$\vec{P} = n\alpha(\vec{E} + \frac{\vec{P}}{3\epsilon_0}) \Rightarrow \vec{P}\left(1 - \frac{n\alpha}{3\epsilon_0}\right) = n\alpha \vec{E}$$

$$\Rightarrow \vec{P} = \frac{n\chi \vec{E}}{(1 - \frac{n\chi}{3\epsilon_0})} \quad \text{And since } \vec{P} = \chi \epsilon_0 \vec{E} \text{ we infer}$$

$$\chi_e = \frac{n\chi/\epsilon_0}{1 - n\chi/3\epsilon_0} \quad \text{Invert this to get}$$

$$\alpha = \frac{\epsilon_0}{n} \left( \frac{\chi_e}{1 + \chi_e/3} \right) \quad \text{and since the dielectric "constant" is defined by } K = 1 + \chi_e \text{ we get}$$

$$\alpha = \frac{3\epsilon_0}{n} \frac{K-1}{K-2}$$

Which works for solids and liquids  
Better than it has any business working given the guys that made it barely knew atoms existed, period.

You get to try numbers in the homework.

### Boundary Conditions

The boundary conditions regarding  $\vec{E}$  still work. At an interface,

$$E_{\text{above}\parallel} = E_{\text{below}\parallel}, \quad E_{\text{above}\perp} - E_{\text{below}\perp} = \sigma/\epsilon_0$$

But  $\sigma$  is  $\sigma_{\text{bound}} + \sigma_{\text{free}}$ , which we don't always know.  
So let's get conditions on  $\vec{D}$ .

If  $\vec{\nabla} \times \vec{P} = 0$  so that  $\vec{\nabla} \times \vec{D} = 0$ , we know  $\oint \vec{D} \cdot d\vec{l} = 0 \quad \forall \text{ paths}$ ,  
so  $D_{\text{above}\parallel} = D_{\text{below}\parallel}$

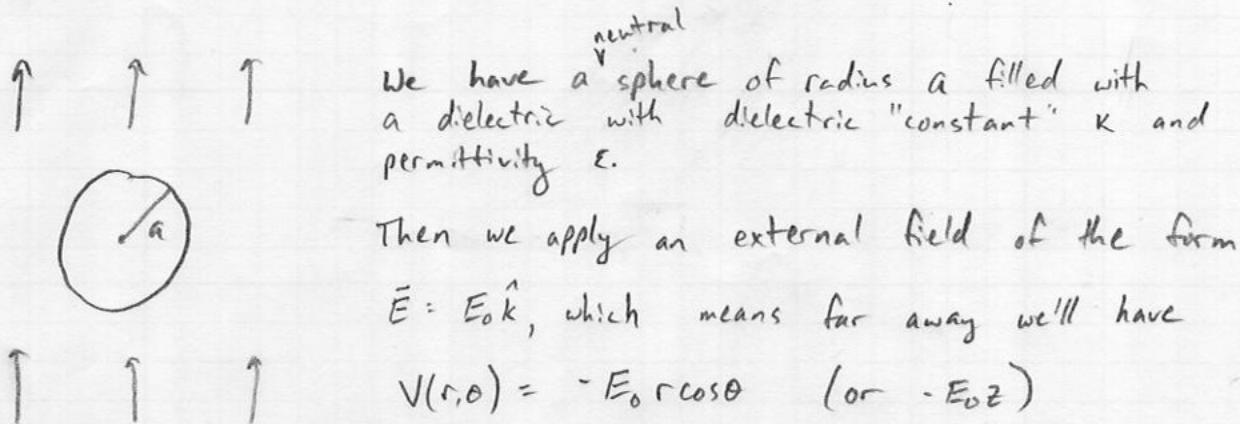
But that's not 100% general.

What is? The condition on  $E_\perp$  was derived from Gauss's Law, and we have  $\oint \vec{D} \cdot d\vec{A} = q_f$ , so

$$D_{\text{above}\perp} - D_{\text{below}\perp} = \sigma_f$$

Where  $\sigma_f$  is usually more accessible than  $\sigma_f + \sigma_b$ .

Example: Spherical solid dielectric in a uniform field



We have a sphere of radius  $a$  filled with a dielectric with dielectric "constant"  $\kappa$  and permittivity  $\epsilon$ .

Then we apply an external field of the form  $\vec{E} = E_0 \hat{k}$ , which means far away we'll have

$$V(r, \theta) = -E_0 r \cos\theta \quad (\text{or } -E_0 z)$$

In spherical coordinates, for systems with no  $\phi$  dependence, the general solution for the potential looks like:

$$V(r, \theta) = \sum_{l=0}^{\infty} (A_l r^l + B_l r^{-l-1}) P_l(\cos\theta)$$

To find the potential everywhere, we need to adapt this to two regions separately. For  $r < a$ , the potential ought to stay finite, so we need all the  $r^{-l-1}$  terms to go away:

$$V(r, \theta)_{\text{in}} = \sum_{l=0}^{\infty} C_l r^l P_l(\cos\theta)$$

Outside, we have to allow for the possibility that  $V$  blows up at large  $r$  on account of the applied field, so we can't exclude all the  $r^l$  terms:

$$V(r, \theta)_{\text{out}} = \sum_{l=0}^{\infty} A_l r^l P_l(\cos\theta) + B_l r^{-l-1} P_l(\cos\theta)$$

To fix our constants, we need boundary conditions. We certainly have

$$V_{\text{out}}(a, \theta) = V_{\text{in}}(a, \theta) \quad (1)$$

We also have

$$V_{\text{out}}(r, \theta) \rightarrow -E_0 r \cos\theta \quad \text{for large } r \quad (2)$$

We technically have  $E_{1\perp} - E_{2\perp} = (\sigma_f + \sigma_b)/\epsilon_0$ ,

but it's not very helpful since we won't know  $\sigma_i$  till after we've solved the problem.

We do, however, know that

$$D_{1\perp} - D_{2\perp} = \sigma_f$$

And  $\sigma_f = 0$ , and  $D = \epsilon E$ , so if we let  $D_1 = D_{\text{out}} = \epsilon_0 E_{\text{out}}$ , we have

$$\epsilon_0 E_{\text{out}\perp} - \epsilon E_{\text{in}\perp} = 0$$

$$\Rightarrow \boxed{\epsilon_0 \frac{\partial V_{\text{out}}}{\partial r} \Big|_{r=a} = \epsilon \frac{\partial V_{\text{in}}}{\partial r} \Big|_{r=a}} \quad (3)$$

Note that since  $\epsilon = \epsilon_0(1 + \chi_e)$ ,  $\epsilon = \epsilon_0$  in vacuum.

Let's use (2) first, since it's a pretty severe constraint. For large  $r$ ,  $V_{\text{out}} \rightarrow -E_0 r \cos\theta$ . Only  $P_1(\cos\theta)$  has that kind of  $\theta$ -dependence, so it must be the case that  $A_1$  is the only nonzero  $A_l$ . One might also guess that holds for the  $B_l$ , but that's not as obvious, so let's leave that be. Then we have

$$V_{\text{out}}(r, \theta) = A_1 r P_1(\cos\theta) + \sum_{l=0}^{\infty} B_l r^{-(l-1)} P_l(\cos\theta)$$

Let's use (1) next. We get:

$$A_1 a P_1(\cos\theta) + \sum_{l=0}^{\infty} B_l a^{-(l-1)} P_l(\cos\theta) = \sum_{l=0}^{\infty} C_l a^l P_l(\cos\theta)$$

This has to hold for all  $\theta$ , so the terms involving each independent  $P_l(\cos\theta)$  need to match individually. This works out slightly differently for  $l=1$  and  $l \neq 1$ . For  $l=1$ :

$$A_1 a P_1(\cos\theta) + B_1 a^{-2} P_1(\cos\theta) = C_1 a P_1(\cos\theta)$$

$$\Rightarrow \boxed{A_1 a + B_1/a^2 = C_1 a}$$

$$\text{For } l \neq 1: \quad B_l a^{-(l-1)} P_l(\cos\theta) = C_l a^l P_l(\cos\theta)$$

$$\Rightarrow \boxed{B_l = C_l a^{2l+1}} \quad (4)$$

Finally, let's use condition (3). r-derivatives of V are easy enough, but again we get slightly different answers for  $l=1$  and  $l \neq 1$ . For  $l=1$ :

$$\epsilon_0 A_1 P_1(\cos\theta) + \epsilon_0 (-2) B_1 \alpha^{-3} P_1(\cos\theta) = \epsilon C_1 P_1(\cos\theta)$$

$$\Rightarrow \boxed{\epsilon_0 A_1 - 2\frac{\epsilon_0 B_1}{\alpha^3} = \epsilon C_1}$$

For  $l \neq 1$  we get:

$$\epsilon_0 (-l-1) \alpha^{(-l-2)} B_l P_l(\cos\theta) = \epsilon l \alpha^{l-1} C_l P_l(\cos\theta)$$

Substituting from (4), we get:

$$\epsilon_0 (-l-1) \alpha^{(-l-2)} C_l \alpha^{2l+1} = \epsilon l \alpha^{l-1} C_l$$

$$\Rightarrow \boxed{[\epsilon_0 (-l-1) \alpha^{l-1}] C_l = [\epsilon l \alpha^{l-1}] C_l}$$

$$\Rightarrow \epsilon_0 (-l-1) C_l = \epsilon l C_l$$

$\epsilon$  is arbitrary, so this has to hold for all  $\epsilon$ . But

in general  $\epsilon_0 (-l-1) \neq \epsilon l$ , so it must be the case that all  $C_l$  (and therefore all  $B_l$ ) are zero if  $l \neq 1$ . Had we guessed that up front, we'd have been right.

So now we're down to our  $l=1$  conditions:

$$A_1 \alpha + B_1 / \alpha^2 = C_1 \alpha$$

$$\epsilon_0 A_1 - 2\frac{\epsilon_0 B_1}{\alpha^3} = C_1 \epsilon$$

We'll tap condition (2) one last time to get

$$A_1 \Gamma P_1(\cos\theta) = -E_0 \Gamma \cos\theta \Rightarrow \boxed{A_1 = -E_0}$$

Moving right along,

$$-E_0 a + B_1/a^2 = a C_1$$

$$-E_0 \epsilon_0 - \frac{2\epsilon_0 B_1}{a^3} = \epsilon C_1$$

so  $C_1 = -E_0 + B_1/a^3$  from the first. Sub into the second:

$$-E_0 \epsilon_0 - \frac{2\epsilon_0 B_1}{a^3} = \epsilon (-E_0 + B_1/a^3)$$

$$\Rightarrow -E_0 \epsilon_0 - 2\epsilon_0 B_1/a^3 = -\epsilon E_0 + \epsilon B_1/a^3$$

$$\Rightarrow (\epsilon - \epsilon_0) E_0 = \left( \frac{\epsilon}{a^3} + 2\frac{\epsilon_0}{a^3} \right) B_1$$

$$\Rightarrow B_1 = \frac{(\epsilon - \epsilon_0)}{\frac{\epsilon}{a^3} + 2\frac{\epsilon_0}{a^3}} = \frac{a^3(\epsilon - \epsilon_0)}{\epsilon + 2\epsilon_0} E_0$$

Here's a fun trick:  $\epsilon = k\epsilon_0$ , so  $\epsilon/\epsilon_0 = k$ . Divide top + bottom by  $\epsilon_0$ :

$$B_1 = \frac{a^3 \left( \frac{\epsilon}{\epsilon_0} - 1 \right)}{\left( \frac{\epsilon}{\epsilon_0} + 2 \right)} E_0$$

$$\Rightarrow B_1 = \frac{(k-1)}{(k+2)} a^3 E_0$$

Back substituting,

$$C_1 = -E_0 + \left( \frac{k-1}{k+2} \right) E_0 = \left[ \frac{k-1}{k+2} - \frac{k+2}{k+2} \right] E_0$$

$$C_1 = -\frac{3}{k+2} E_0$$

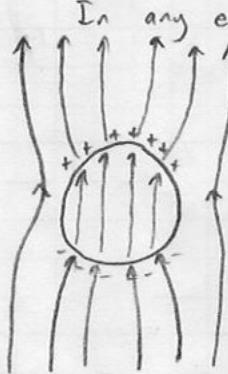
Yielding, finally,

$$V_{in}(r, \theta) = -\frac{3E_0}{K+2} r \cos\theta = -\frac{3E_0}{K+2} z$$

$$V_{out}(r, \theta) = -E_0 r \cos\theta + \frac{E_0 a^3}{r^2} \frac{(K-1)}{(K+2)} \cos\theta$$

We can see  $\vec{E}_{inside} = -\frac{dV_{in}}{dr} \hat{r} = \frac{3E_0}{K+2} \hat{r}$  Note this is basically  $\vec{E}_{applied}$ , but a little weaker since  $K > 1$ . As  $K \rightarrow \infty$   $E_{inside} \rightarrow 0$  (conducting limit)

In any event the polarized charges cancel out some of the applied field



Let's get  $\vec{P}$  so we can check  $\sigma_b$  and  $\sigma_f$ .

$\vec{P}$  exists inside the sphere, so

$$\vec{P} = \chi_e \epsilon_0 \left( \frac{3E_0}{K+2} \right) \vec{r} \quad \text{and} \quad \chi_e = K-1$$

$$\vec{P} = \frac{(K-1)}{(K+2)} 3E_0 \epsilon_0 \hat{r}$$

And  $\rho_b = -\nabla \cdot \vec{P} = 0$ , which we expect.

$$\sigma_b = \hat{n} \cdot \vec{P} = \hat{r} \cdot \vec{P} (\text{blar}) = 3E_0 \epsilon_0 \frac{(K-1)}{(K+2)} \cos\theta$$

Which has the kind of  $\theta$  dependence we expect.