

- 1) HM problem 9-21.
- 2) HM problem 9-22.
- 3) When the dipole radiates because it is driven by an external wave, the incident radiation is *scattered*. We can use this description of scattering when the scattering object is much less than the wavelength.

- a. A forced, damped oscillator has a position dependence:

$$x(t) = \frac{-eE_0}{m_e} \frac{e^{-i\omega t}}{(\omega_0^2 - \omega^2) - i\nu\omega},$$

where  $\omega_0$  is the resonance frequency, and  $\nu$  is the radiative damping rate for a particular resonance (see class notes for the radiation damping of a freely oscillating charge). Calculate the total radiated power, using the Larmor equation (eq. 8-89 in HM). To calculate the acceleration, do the time derivatives, then take the real part, that is  $\langle a^2 \rangle = \langle (\text{Re}[\ddot{x}])^2 \rangle$ .

- b. The scattering cross-section  $\sigma$  is calculated by dividing the radiated power by the incident intensity:  $\sigma = P_{\text{avg}} / I_{\text{inc}}$ . Show that the scattering cross-section is given by:

$$\sigma = \frac{8\pi r_e^2}{3} \frac{\omega^4}{(\omega_0^2 - \omega^2)^2 + (\nu\omega)^2}, \text{ where the classical electron radius is } r_e = e^2 / m_e c^2.$$

- c. Show that in the low frequency limit, i.e.  $\omega \ll \omega_0$ ,  $\nu \ll \omega_0$ ,  $\sigma = \frac{8\pi r_e^2}{3} \left( \frac{\omega}{\omega_0} \right)^4$ .

In this case, we have Rayleigh scattering, with the characteristic  $\omega^4$  dependence. (One reason why the sky is blue.)

- d. Show that in the high frequency limit, i.e.  $\omega \gg \omega_0$ ,  $\sigma = \frac{8\pi r_e^2}{3}$ . This is the limit of Thomson scattering, the scattering from a free electron. When the frequency of the incident light is above all resonances (for example soft x-rays incident on atoms), the electrons behave as if they are free.
- e. Make a plot of  $\text{Log}_{10}[\sigma]$  vs  $\omega/\omega_0$  for a range near  $0 < \omega/\omega_0 < 3$ , for two damping rates,  $\nu = 0.1 \omega_0$  and  $\nu = 0.02 \omega_0$ .

- 4) Consider a collection of free electrons that have a thermal velocity distribution (Boltzmann). The Thomson-scattered light will be Doppler shifted ( $\Delta\omega = \pm\omega_0 v / c$ , where  $\omega_0$  is the angular frequency of the incident light).

- a. Show that the spectrum of the scattered light is Gaussian in profile.
- b. Derive an expression for the full-width at half maximum (FWHM) of the spectrum and the temperature of the plasma. What mean thermal energy (kT) in eV would correspond to a measured width of  $\Delta\omega / \omega_0 = 5 \times 10^{-3}$ ?

- 5) HM problem 10.4.