

# Interaction of light with atoms: Rate equations

Rate equations: connect wave propagation to atom dynamics

Einstein A and B coefficients

# Transition rates and rate equations

- We have been looking from the point of view of the photons. What about the atoms?
  - Absorption of a photon induces a transition from level 1 to 2.

$$\frac{dN_1}{dt} = -N_1 W_{12} \quad \frac{dN_2}{dt} = N_1 W_{12} = -\frac{dN_1}{dt}$$

- The absorption rate  $W$  must depend on the intensity and the incident frequency. We'll represent this by the spectral energy density.
- For light at a *specific* frequency, define
$$W_{12} = B_{12} \rho(\nu_0) \quad B_{12} = \text{Einstein "B" coefficient}$$
- Will generalize later for broadband light

# Spontaneous emission

- An atom in an excited state can decay to another level through radiation = spontaneous emission

$$\frac{dN_2}{dt} = -N_2 A_{21} \rightarrow N_2(t) = N_2(0) e^{-A_{21}t}$$

Lifetime of state:  
 $\tau_2 = 1 / A_{21}$

- If there are multiple destination states, rates add.  
Total decay out of level  $i$  :

$$\frac{dN_i}{dt} = -N_i \sum_j A_{ij}$$

Lifetime of state:  
 $\tau_i = 1 / \sum_j A_{ij}$

- Note this type of process is independent of any incident light.

# Einstein's treatment of emission and absorption

- Based on thermodynamic principles, Einstein predicted the existence of stimulated emission.
- First suppose we have *only* absorption and spontaneous emission.
- Rate equations for a two-level system (no SE):

$$\frac{dN_1}{dt} = -N_1 B_{12}\rho(\nu) + N_2 A_{21} \quad \frac{dN_2}{dt} = +N_1 B_{12}\rho(\nu) - N_2 A_{21}$$

- In equilibrium with the field, no net change in population densities

$$0 = -N_1^e B_{12}\rho(\nu) + N_2^e A_{21} \rightarrow \frac{N_2^e}{N_1^e} = \frac{B_{12}\rho(\nu)}{A_{21}}$$

# Thermal equilibrium with BB field

- An atom that is in thermal equilibrium has populations that follow the Boltzmann distribution:

$$\frac{N_2^e}{N_1^e} = \frac{g_2}{g_1} e^{-hv_{21}/k_B T} = \frac{B_{12}\rho(\nu)}{A_{21}} \rightarrow \rho(\nu) = \frac{A_{21}}{B_{12}} \frac{g_2}{g_1} e^{-hv_{21}/k_B T}$$

- A field in thermal equilibrium should have the blackbody spectral energy density

$$\rho_{BB}(\nu) = 8\pi \frac{\nu^2}{c^3} \frac{h\nu}{e^{h\nu/k_B T} - 1}$$

- What we have is ok in the high frequency limit, but not fully consistent with the BB curve.

## Including stimulated emission

- Things make more sense if we allow for another route for transition from 2 to 1
  - Add stimulated emission to rate equations:

$$\frac{dN_1}{dt} = -N_1 B_{12}\rho(\nu) + N_2 B_{21}\rho(\nu) + N_2 A_{21} \quad \frac{dN_2}{dt} = -\frac{dN_1}{dt}$$

- Equilibrium:  $d/dt = 0$

$$0 = -N_1^e B_{12}\rho(\nu) + N_2^e B_{21}\rho(\nu) + N_2^e A_{21} \rightarrow \frac{N_2^e}{N_1^e} = \frac{B_{12}\rho(\nu)}{A_{21} + B_{21}\rho(\nu)}$$

$$\frac{N_2^e}{N_1^e} = \frac{g_2}{g_1} e^{-h\nu_{21}/k_B T} = \frac{B_{12}\rho(\nu)}{A_{21} + B_{21}\rho(\nu)}$$

# Equilibrium spectral energy density

- Solve for the equilibrium spectral energy density

$$\frac{N_2^e}{N_1^e} = \frac{g_2}{g_1} e^{-h\nu_{21}/k_B T} = \frac{B_{12}\rho(\nu)}{A_{21} + B_{21}\rho(\nu)}$$

$$\frac{g_2}{g_1} e^{-h\nu_{21}/k_B T} (A_{21} + B_{21}\rho(\nu)) = B_{12}\rho(\nu)$$

$$\rho(\nu) = \frac{A_{21}}{B_{12} \frac{g_1}{g_2} e^{h\nu_{21}/k_B T} - B_{21}}$$

- This looks similar in form to the blackbody relation

# Einstein's relations between A and B coefficients

- If both the atoms and BB cavity are in thermal equilibrium, the  $\rho(\nu)$ 's that satisfy that constraint must be the same

$$\rho_{BB}(\nu) = 8\pi \frac{\nu^2}{c^3} \frac{h\nu}{e^{h\nu/k_B T} - 1} \qquad \rho(\nu) = \frac{A_{21}}{B_{12} \frac{g_1}{g_2} e^{h\nu_{21}/k_B T} - B_{21}}$$

- The two forms will have the same structure if

$$B_{12} \frac{g_1}{g_2} = B_{21} \rightarrow \rho(\nu) = \frac{A_{21}}{B_{21} (e^{h\nu_{21}/k_B T} - 1)}$$

- So the processes of absorption and stimulated emission are linked.
- Finally, for  $\rho_{BB}(\nu) = \rho(\nu)$

$$A_{21} = \frac{8\pi h\nu^3}{c^3} B_{21}$$



# Physical significance of A/B

- Dimensionally,  $B_{21}\rho$  gives a rate, so in the relation between A and B,

$$A_{21} = \frac{8\pi h\nu^3}{c^3} B_{21}$$

$\rho(\nu) = \frac{8\pi h\nu^3}{c^3}$  is a type of spectral energy density.

In QED, the E and B energy densities are quantized, and the quanta are the photons.

$\rho(\nu) = \frac{8\pi h\nu^3}{c^3}$  is effectively the spectral energy density of the vacuum fluctuations of the field.

## Connect intensity changes to atomic rates

- In a volume  $V$ , absorbed power is  $\frac{dP_a}{dV} = W_{12} N_1 h\nu$
- For a beam with area  $A$ ,  $\frac{dP_a}{dV} = \frac{1}{A} \frac{dP}{dz} = -\frac{dI}{dz}$
- Intensity and energy density are related:  $\rho c = I$

$$\frac{dP_a}{dV} = -\frac{dI}{dz} = B_{12} \rho N_1 h\nu$$

$$\frac{dI}{dz} = -I N_1 \frac{B_{12} h\nu}{c} = -I N_1 \sigma_{12}$$

$$\sigma_{12} = \frac{B_{12} h\nu}{c}$$

Compare to earlier generalization to account for lineshape of absorption, and bandwidth of source.

Note that the mean free path of photons in the medium is  $1/\alpha$

# Optical gain

- With population in both levels 1 and 2,

$$\frac{dI}{dz} = I \left( N_2 B_{21} - N_1 B_{12} \right) \frac{h\nu_{21}}{c} \quad B_{12} \frac{g_1}{g_2} = B_{21}$$

$$\frac{dI}{dz} = I \left( N_2 - N_1 \frac{g_2}{g_1} \right) \frac{B_{21} h\nu_{21}}{c} = I N_{inv} \sigma_{21}$$

Inversion  
density

Gain  
cross-  
section

$$I(z) = I_0 e^{gz}$$

$g$ : gain coefficient =  $N_{inv} \sigma_{21}$   
(opposite sign from absorption coefficient)

For an amplifier of length  $L$ ,

$$I(L) = I_0 e^{gL} = I_0 G_0$$

$G_0$ : small signal single-pass gain