

You may use other resources for these (other texts, web), but cite what you use.

- 1) The following are useful for making estimates – you should memorize these relations:
 - a. Wavelength of a photon that has 1eV in energy
 - b. Energy (in eV) of a photon with wavelength of 1 μ m.
 - c. Energy (in eV) for kT at room temperature (300K).
 - d. Show that for light, $\frac{\Delta\lambda}{\lambda_0} = \frac{\Delta\omega}{\omega_0} = \frac{\Delta\nu}{\nu_0} = \frac{\Delta\sigma}{\sigma_0}$, where the Δ 's correspond to bandwidth, and the denominators correspond to the central carrier frequency (or wavelength...). The symbols, in order are: wavelength, angular frequency, frequency, and wavenumber. In these expressions, you can drop any sign you get in the derivation.
 - e. These relations are useful to convert spectral bandwidth into different units. As an example, calculate the bandwidth in nm ($\Delta\lambda$) for a signal with $\Delta\nu = 4 \times 10^{13}$ Hz for the two cases, where the central wavelength is 800nm and 200nm. (Background: the bandwidth required for a laser pulse of a given duration is inversely proportional to $\Delta\nu$, independent of the central frequency or wavelength.)
- 2) Hooker & Webb (H&W) problem 1.2
- 3) H&W problem 1.3. For part b, consider a photon moving at a velocity c an angle to a plane with a hole of area A . The projection of the velocity perpendicular to the plane is lower, as is the hole area. Average over the angles to get the factor.
- 4) H&W problem 2.1
- 5) H&W problem 2.2