

5-3. In order to remove a particle from the surface of the Earth and transport it infinitely far away, the initial kinetic energy must equal the work required to move the particle from $r = R_e$ to $r = \infty$ against the attractive gravitational force:

$$\int_{R_e}^{\infty} G \frac{M_e m}{r^2} dr = \frac{1}{2} m v_0^2 \quad (1)$$

where M_e and R_e are the mass and the radius of the Earth, respectively, and v_0 is the initial velocity of the particle at $r = R_e$.

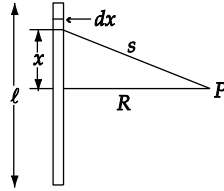
Solving (1), we have the expression for v_0 :

$$v_0 = \sqrt{\frac{2G M_e}{R_e}} \quad (2)$$

Substituting $G = 6.67 \times 10^{-11} \text{ m}^3/\text{kg} \cdot \text{s}^2$, $M_e = 5.98 \times 10^{24} \text{ kg}$, $R_e = 6.38 \times 10^6 \text{ m}$, we have

$$\boxed{v_0 \cong 11.2 \text{ km/sec}} \quad (3)$$

5-7.



The contribution to the potential at P from a small line element is

$$d\Phi = -G \frac{\rho_l}{s} dx \quad (1)$$

where $\rho_l = \frac{M}{l}$ is the linear mass density. Integrating over the whole rod, we find the potential

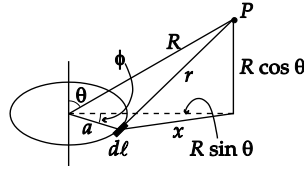
$$\Phi = -G \frac{M}{l} \int_{-\ell/2}^{\ell/2} \frac{1}{\sqrt{x^2 + R^2}} dx \quad (2)$$

Using Eq. (E.6), Appendix E, we have

$$\Phi = -G \frac{M}{l} \ln \left[x + \sqrt{x^2 + R^2} \right]_{-\ell/2}^{\ell/2} = -\frac{GM}{l} \ln \left[\frac{\frac{\ell}{2} + \sqrt{\frac{\ell^2}{4} + R^2}}{-\frac{\ell}{2} + \sqrt{\frac{\ell^2}{4} + R^2}} \right]$$

$$\boxed{\Phi = -\frac{GM}{l} \ln \left[\frac{\sqrt{\ell^2 + 4R^2} + \ell}{\sqrt{\ell^2 + 4R^2} - \ell} \right]} \quad (3)$$

5-10.



Using the relations

$$x = \sqrt{(R \sin \theta)^2 + a^2 - 2aR \sin \theta \cos \phi} \quad (1)$$

$$r = \sqrt{x^2 + R^2 \cos^2 \theta} = \sqrt{R^2 + a^2 - 2aR \sin \theta \cos \phi} \quad (2)$$

$$\rho_\ell = \frac{M}{2\pi a} \text{ (the linear mass density),} \quad (3)$$

the potential is expressed by

$$\Phi = -G \int \frac{\rho_\ell dl}{r} = -\frac{GM}{2\pi R} \int_0^{2\pi} \frac{d\phi}{\sqrt{1 - \left[2 \frac{a}{R} \sin \theta \cos \phi - \frac{a^2}{R^2} \right]}} \quad (4)$$

If we expand the integrand and neglect terms of order $(a/R)^3$ and higher, we have

$$\left[1 - \left[2 \frac{a}{R} \sin \theta \cos \phi - \frac{a^2}{R^2} \right] \right]^{-1/2} \cong 1 + \frac{a}{R} \sin \theta \cos \phi - \frac{1}{2} \frac{a^2}{R^2} + \frac{3}{2} \frac{a^2}{R^2} \sin^2 \theta \cos^2 \phi \quad (5)$$

Then, (4) becomes

$$\Phi \cong -\frac{GM}{2\pi R} \left[2\pi - \frac{1}{2} \frac{a^2}{R^2} 2\pi + \frac{3}{2} \frac{a^2}{R^2} \pi \sin^2 \theta \right]$$

Thus,

$$\boxed{\Phi(R) \cong -\frac{GM}{R} \left[1 - \frac{1}{2} \frac{a^2}{R^2} \left[1 - \frac{3}{2} \sin^2 \theta \right] \right]} \quad (6)$$

5-14. Think of assembling the sphere a shell at a time ($r = 0$ to $r = R$).

For a shell of radius r , the incremental energy is $dU = dm \phi$ where ϕ is the potential due to the mass already assembled, and dm is the mass of the shell.

So

$$dm = \rho 4\pi r^2 dr = \left[\frac{3M}{4\pi R^3} \right] 4\pi r^2 dr = \frac{3Mr^2 dr}{R^3}$$

$$\phi = -\frac{Gm}{r} \text{ where } m = M \frac{r^3}{R^3}$$

So

$$\begin{aligned}U &= \int du \\&= \int_{r=0}^R \left[\frac{3Mr^2 dr}{R^3} \right] \left[-\frac{GMr^2}{R^3} \right] \\&= -\frac{3GM^2}{R^6} \int_0^R r^4 dr \\&\boxed{U = -\frac{3}{5} \frac{GM^2}{R}}\end{aligned}$$