

2 beam interference

- measure intensity; but fields add  
intensity for arbitrary polarization:

$$I = \frac{1}{2} \epsilon_0 c n \vec{E} \cdot \vec{E}^* = \frac{1}{2} \epsilon_0 c n (E_x E_x^* + E_y E_y^* + E_z E_z^*)$$

combine two waves

different polarization

$$I = \frac{1}{2} \epsilon_0 c n (E_1 \hat{x} + E_2 \hat{y}) \cdot (E_1^* \hat{x} + E_2^* \hat{y})$$

$$= \frac{1}{2} \epsilon_0 c n (|E_1|^2 + |E_2|^2) = I_1 + I_2$$

since intensities add no interference (no cross-terms)

same polarization

$$I = \frac{1}{2} \epsilon_0 c n ((E_1 + E_2) \hat{x}) \cdot (E_1^* + E_2^*) \hat{x}$$

$$= I_1 + \frac{1}{2} \epsilon_0 c n \cdot 2 \operatorname{Re}(E_1 E_2^*) + I_2$$

cross term shows  
interference.

In general, orthogonal states of pol. don't interfere.

$\hat{x}, \hat{y}$  or R, L circular.

Combine two equal plane waves w/ phase difference  $\delta$

$$E_{tot} = E_0 + E_0 e^{i\delta}$$

$$\begin{aligned} I_{tot} &\propto |E_{tot}|^2 = |E_0|^2 |1 + e^{i\delta}|^2 \\ &= |E_0|^2 (1 + 1 + e^{i\delta} + e^{-i\delta}) \\ &= |E_0|^2 (2 + 2\cos\delta) \end{aligned}$$

when phase difference  $\delta = m\pi$ ,

$$|E_{tot}|^2 = 4|E_0|^2$$

constructive interference

when phase diff =  $\delta = m\pi + \pi/2$

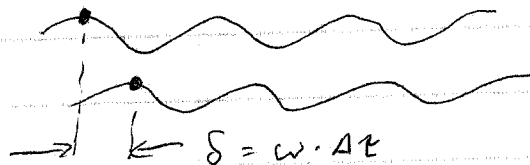
$$|E_{tot}|^2 = 0$$

destructive intenf.

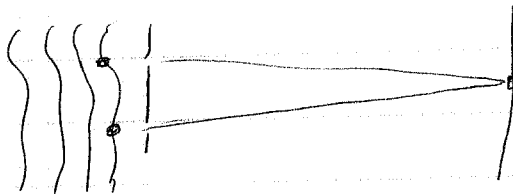
$\delta$  = phase difference, <sup>optical</sup> path diff., reflection phase shifts  
 =  $\omega(t_1 - t_2)$  time diff  
 =  $k_0(n_1 z_1 - n_2 z_2)$  optical path diff.

Ultimately, we compare a wave with another wave,

- amplitude split: compare wave w/ copy at later time




- wavefront split: compare diff portions of wavefront  
 > either same time, or diff't time



Temporal coherence time  $\tau_c$

incoherence is random phase

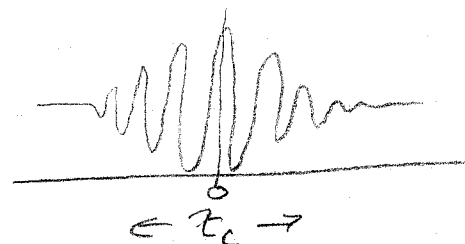
- single atom 

collisions interrupt phase

- multi-atom:  $\therefore$  initial phase is all random.

net effect: light wave keeps its coherence for an average time  $\tau_c$

Measure  $\tau_c$ :  $I(x)$

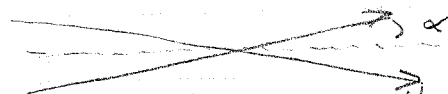


Fringes wash out

reduced contrast or visibility.

$$\frac{I_{\max} - I_{\min}}{I_{\max} + I_{\min}}$$

Misaligned Michelson: tilt mirrors



as before,  $E = E_0 e^{ik_z z} (e^{ik_x x} + e^{-ik_x x})$

$$k_x = k_0 \sin \alpha$$

$$|E|^2 \sim |E_0|^2 \cos^2 k_x x$$

if arm lengths are initially equal, space interference pattern follows time pattern.

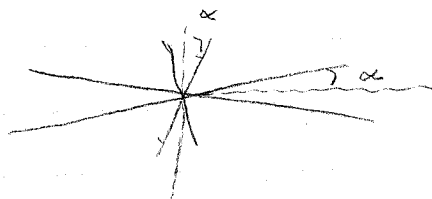
$\alpha$  very small

$$\lambda \text{ period} \sim \frac{2\pi}{2k_0 \sin \alpha}$$

$$\Delta x = \frac{\lambda}{2 \sin \alpha}$$

for  $\Delta x \sim 100 \mu\text{m}$   $\lambda \sim 0.5 \mu\text{m}$

$$\alpha \sim \frac{\lambda}{2\Delta x} \approx 2.5 \text{ mrad} = 0.14^\circ$$



$$\tau(x) = \pm \frac{x \tan \alpha}{c}$$

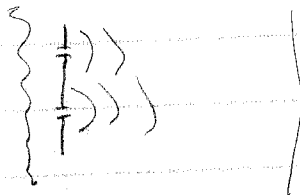
$$E \sim E_0 (e^{ik_x \tan \alpha} + e^{-ik_x \tan \alpha})$$

$$\sim 2E_0 \cos(k_x \tan \alpha)$$

$$\omega \Delta \tau(x) = \pm \frac{\omega}{c} x \tan \alpha$$

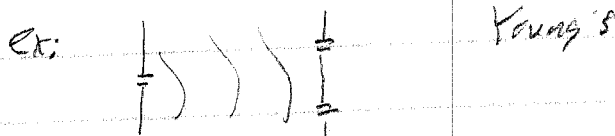
## Spatial coherence length

sample wavefront across beam, compare  
 measure: vary separation of two point sources



extended sources - lamps, sun  
 have very small spatial coherence lengths

Either sample small portion of wavefront.



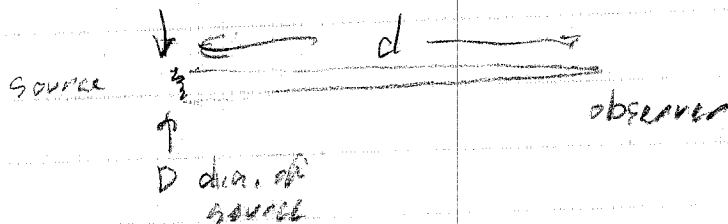
or spectrometer w/ narrow entrance slit.

Or more very far away: starlight has high spatial coh.  
 - stellar interferometry: can measure size of star.

## Van Cittert-Zernike theorem

spatial coherence area is

$$A_c = \frac{D^2 \lambda^2}{\pi d^2} = \frac{\lambda^2}{\Delta \Omega}$$



wavefronts smooth out w/ distance.