

Working with spectral lineshapes

- For atomic system, replace Dirac delta with transition lineshape

$$\int g(\nu - \nu_0) d\nu = 1$$

- Lorentzian lineshape (radiative, collisional broadening)

$$\delta(\nu - \nu_0) \rightarrow g_L(\nu - \nu_0) = \frac{2}{\pi \Delta\nu_0} \frac{1}{1 + \left(\frac{2(\nu - \nu_0)}{\Delta\nu_0} \right)^2}$$

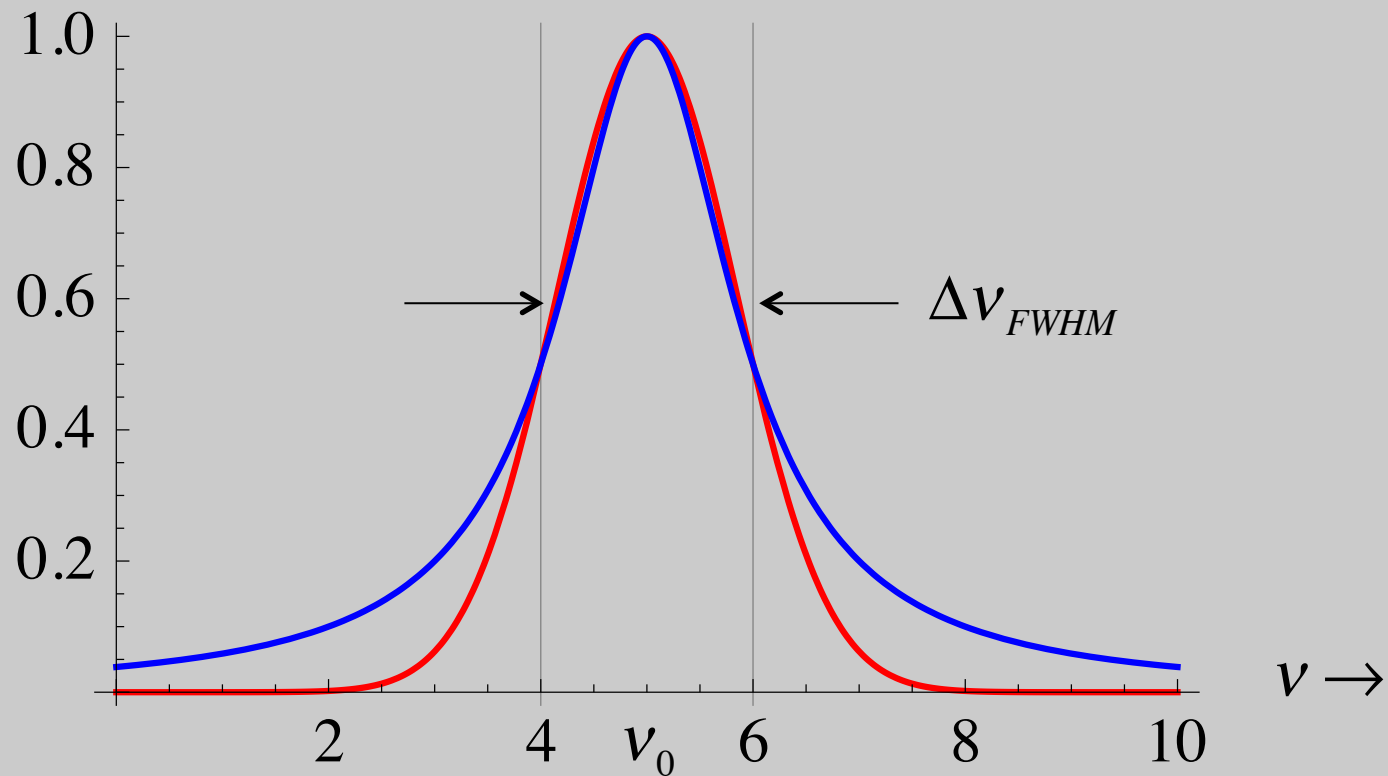
$\Delta\nu_0$ FWHM

- Doppler broadened (Gaussian) lineshape

$$\delta(\nu - \nu_0) \rightarrow g_G^*(\nu - \nu_0) = \frac{2}{\Delta\nu_0^*} \sqrt{\frac{\ln 2}{\pi}} \exp \left\{ -4 \ln 2 \frac{(\nu - \nu_0)^2}{\Delta\nu_0^{*2}} \right\}$$

Lorentzian vs Gaussian lineshapes

- Lorentzian is much broader in spectral wings than Gaussian



Natural broadening

- Radiative broadening comes from the spontaneous emission lifetime of the state: energy-time uncertainty
- Fourier transforms

- Forward: FT $F(\omega) = \int_{-\infty}^{\infty} f(t) e^{i\omega t} dt$

- Inverse: FT⁻¹ $f(t) = \frac{1}{2\pi} \int_{-\infty}^{\infty} F(\omega) e^{-i\omega t} d\omega$

- Suppose exponential, oscillating decay in time domain

$$f(t) = \begin{cases} e^{-\gamma t} e^{-i\omega_0 t} & \text{for } t \geq 0 \\ 0 & \text{for } t < 0 \end{cases}$$

$$F(\omega) = \int_0^{\infty} e^{-\gamma t - i\omega_0 t} e^{i\omega t} dt = \frac{e^{(-\gamma + i(\omega - \omega_0))t}}{-\gamma + i(\omega - \omega_0)} \Big|_0^{\infty} = \frac{1}{\gamma - i(\omega - \omega_0)}$$

Complex Lorentzian

Lorentzian lineshape

- Complex Lorentzian separated into Re and Im

$$\frac{1}{\gamma - i(\omega - \omega_0)} = \frac{\gamma}{(\omega - \omega_0)^2 + \gamma^2} + i \frac{(\omega - \omega_0)}{(\omega - \omega_0)^2 + \gamma^2}$$

– Real part corresponds to absorption effects

- Normalize

$$c \int \frac{\gamma}{(\omega - \omega_0)^2 + \gamma^2} d\omega = c \gamma \frac{\pi}{\gamma} = 1 \quad \rightarrow \quad g_L(\omega - \omega_0) = \frac{\gamma / \pi}{(\omega - \omega_0)^2 + \gamma^2}$$

- Convert ω to ν

$$c \int \frac{\gamma}{4\pi^2 (\nu - \nu_0)^2 + \gamma^2} d\nu = c \gamma \frac{1}{2\gamma} = 1$$

$$\rightarrow g_L(\nu - \nu_0) = \frac{2}{\gamma} \left[1 + \left(\frac{2(\nu - \nu_0)}{\gamma / \pi} \right)^2 \right]^{-1} = \frac{2}{\pi \Delta\nu_0} \left[1 + \left(\frac{2(\nu - \nu_0)}{\Delta\nu_0} \right)^2 \right]^{-1}$$

Collisional broadening

- Elastic collisions don't cause transition, but interrupt the phase

- Timescales:

- Period of EM cycle much less than radiative lifetime $\frac{2\pi}{\omega_0} \ll \tau$

- Avg time btw collisions < lifetime $\tau_c < \tau$

- Duration of a collision \ll time btw coll, lifetime $\Delta\tau_c \ll \tau_c, \tau$

- Calculation:

- FT over time 0 to τ_1 to get lineshape for a specific oscillation length
- Average over probability of a given time between collisions:

$$P(\tau_1)d\tau_1 = \frac{1}{\tau_c} e^{-\tau_1/\tau_c} d\tau_1$$

Result:

Lorentzian shape with new width

$$\Delta\nu = \gamma / 2\pi + 1 / \pi \tau_c$$

Doppler broadening

- From relative velocity of atom to input beam, Doppler shift:

$$v'_0 = \frac{v_0}{1 - v_z / c} \quad \text{Beam propagating in z direction}$$

- Each atom in distribution is shifted according to its velocity

- Boltzmann distribution

$$P(v_z) \sim \exp\left[-\frac{1}{2} M v_z^2 / k_B T\right]$$

- Average over distribution to get effective lineshape:

$$g^*(\nu - \nu_0) = \frac{1}{\nu_0} \left(\frac{Mc^2}{2\pi k_B T} \right)^{1/2} \exp\left\{ \frac{Mc^2}{2k_B T} \frac{(\nu - \nu_0)^2}{\nu_0^2} \right\} \quad \text{Gaussian lineshape}$$

FWHM: $\Delta\nu_0^* = 2\nu_0 \left[\frac{2k_B T \ln 2}{Mc^2} \right]^{1/2}$

Doppler broadening in HeNe lasers

$$\Delta\nu_0^* = 2\nu_0 \left[\frac{2k_B T \ln 2}{Mc^2} \right]^{1/2}$$

$$\lambda_0 = 632.8 \text{ nm}$$

$$\nu_0 = 4.74 \times 10^{14} \text{ s}^{-1}$$

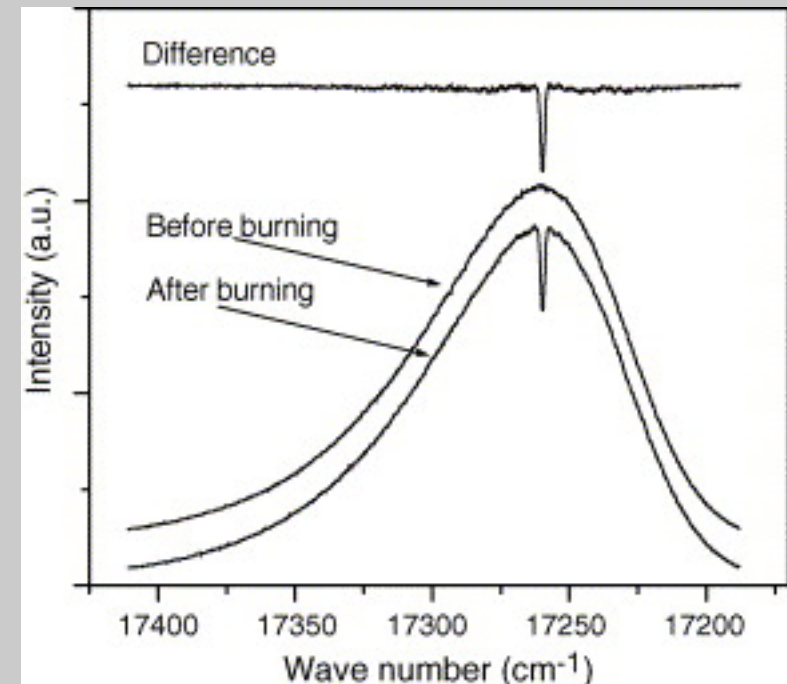
$$M = 20.12 \text{ amu} = 3.34 \times 10^{-26} \text{ kg} \quad \text{For Neon}$$

$$k_B T = 1/40 \text{ eV} = 4 \times 10^{-21} \text{ J}$$

$$\Delta\nu_0^* = 1.55 \text{ GHz}$$

Inhomogeneous vs homogeneous broadening

- Homogeneous broadening:
every atom is broadened by same shape
 - Radiative, collisional, phonon
 - All atoms participate in absorption or gain
- Inhomogeneous broadening:
 - Doppler broadening
 - Absorption or gain only by atoms in resonance
 - Leads to “spectral hole burning”



Fermi's golden rule generalized

- To account for the transition lineshape:

$$W_{12}(\nu) = \frac{2\pi^2}{3n^2\epsilon_0 h^2} |\mu_{12}|^2 \rho_\nu \delta(\nu - \nu_0) \rightarrow \frac{2\pi^2}{3n^2\epsilon_0 h^2} |\mu_{21}|^2 \rho_\nu g(\nu - \nu_0)$$

- Example: narrow linewidth laser incident on atom

$$\rho_\nu = \frac{n}{c} I \delta(\nu - \nu_L)$$

- Total transition rate:

$$W_{12} = \int \frac{2\pi^2}{3n^2\epsilon_0 h^2} |\mu_{12}|^2 \frac{n}{c} I \delta(\nu - \nu_L) g(\nu - \nu_0) d\nu$$

$$W_{12}(\nu_L) = \frac{2\pi^2}{3n\epsilon_0 c h^2} |\mu_{12}|^2 I g(\nu_L - \nu_0)$$

Cross sections

- It is inconvenient to carry around all these constants and to use the dipole moments
 - Use values that connect to what we can measure

- Consider a beam passing through a gas of atoms with number density N_t (atoms/unit volume)

- Power absorbed/unit volume:
$$\frac{dP_a}{dV} = W_{12} N_t h\nu$$

- Photon flux (photons/area/sec)
$$F = \frac{I}{h\nu}$$

- Power in beam:
$$P = I A = h\nu F A$$

- Evolution of flux:
$$\frac{dF}{dz} = \frac{1}{h\nu} \frac{1}{A} \frac{dP}{dz} = -\frac{1}{h\nu} \frac{dP_a}{dV} = -W_{12} N_t$$

Cross sections

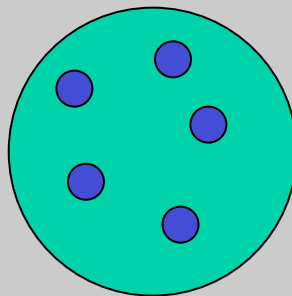
- Flux decreases as beam propagates in medium

$$\frac{dF}{dz} = -W_{12}N_t$$

- Define total absorption cross-section: $\sigma_{12} = \frac{W_{12}}{F}$

- So that: $\frac{dF}{dz} = -N_t\sigma_{12} \rightarrow F(z) = F_0 e^{-N_t\sigma_{12}z}$

- Physically, the cross-section is an effective area of the atom. In a low density gas, the beam of photons sees a collection of spheres:



$$\frac{dF}{F} = -N_t A dz \frac{\sigma_{12}}{A}$$

Frequency-dependent cross section

- Total cross section is obtained by integrating over lineshape (h for homogeneous):

$$\sigma_{12} = \int \sigma_h(\nu) d\nu$$

- Suppose we have a narrowband laser beam with a frequency that we can tune

$$W_{12}(\nu_L) = \frac{2\pi^2}{3n\epsilon_0 ch^2} |\mu_{12}|^2 I g(\nu_L - \nu_0)$$

$$\sigma_h(\nu_L) = \frac{W_{12}(\nu_L)}{F} = \frac{W_{12}(\nu_L)}{I/h\nu_L} = \frac{h\nu_L}{I} \frac{2\pi^2}{3n\epsilon_0 ch^2} |\mu_{12}|^2 I g(\nu_L - \nu_0)$$

$$\sigma_h(\nu_L) = \frac{2\pi^2}{3n\epsilon_0 ch} |\mu_{12}|^2 \nu_L g(\nu_L - \nu_0)$$