

- 1) HM problem 9.4
- 2) HM problem 9-12.
- 3) HM problem 9-16. This impedance-matching problem is not specific to antennas – it works for any load resistance.
- 4) HM problem 9-17: You can consider the rotating dipole as a superposition of two linear dipoles oriented along the x and y axes.
- 5) When the dipole radiates because it is driven by an external wave, the incident radiation is *scattered*. We can use this description of scattering when the scattering object is much less than the wavelength.

a. A forced, damped oscillator has a position dependence:

$$x(t) = \frac{-eE_0}{m_e} \frac{e^{-i\omega t}}{(\omega_0^2 - \omega^2) - i\nu\omega},$$

where ω_0 is the resonance frequency, and ν is the radiative damping rate for a particular resonance (see class notes for the radiation damping of a freely oscillating charge). Calculate the total radiated power, using the Larmor equation (eq. 8-89 in HM). To calculate the acceleration, do the time derivatives, then take the real part, that is $\langle a^2 \rangle = \langle (\text{Re}[\ddot{x}])^2 \rangle$.

b. The scattering cross-section σ is calculated by dividing the radiated power by the incident intensity: $\sigma = P_{\text{avg}} / I_{\text{inc}}$. Show that the scattering cross-section is given by:

$$\sigma = \frac{8\pi r_e^2}{3} \frac{\omega^4}{(\omega_0^2 - \omega^2)^2 + (\nu\omega)^2}, \text{ where the classical electron radius is } r_e = e^2 / m_e c^2.$$

c. Show that in the low frequency limit, i.e. $\omega \ll \omega_0$, $\nu \ll \omega_0$, $\sigma = \frac{8\pi r_e^2}{3} \left(\frac{\omega}{\omega_0} \right)^4$.

In this case, we have Rayleigh scattering, with the characteristic ω^4 dependence. (One reason why the sky is blue.)

d. Show that in the high frequency limit, i.e. $\omega \gg \omega_0$, $\sigma = \frac{8\pi r_e^2}{3}$. This is the limit of Thomson scattering, the scattering from a free electron.

e. Make a plot of $\text{Log}_{10}[\sigma]$ vs ω/ω_0 for a range near $0 < \omega/\omega_0 < 3$, for two damping rates, $\nu = 0.1 \omega_0$ and $\nu = 0.02 \omega_0$.